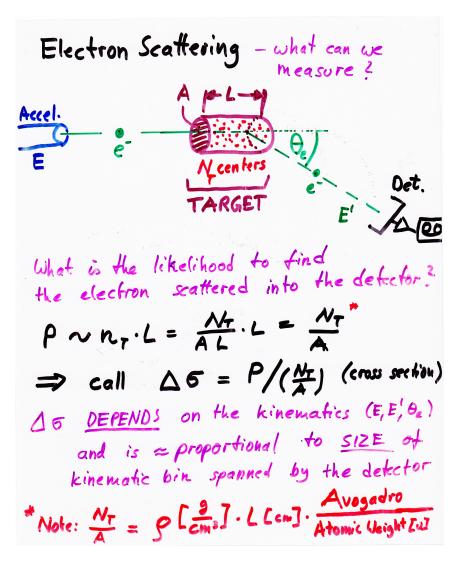
OBSERVABLES



Count rate (L = luminosity): $\dot{N} = P \cdot \dot{n}_{el} = \Delta \sigma \cdot \frac{N_T}{A} \dot{n}_{el} = \Delta \sigma \cdot \frac{N_T}{A} \frac{I}{e} = \Delta \sigma \cdot L$ In general: $\Delta \sigma = \Delta \sigma (E' ... E' + \Delta E', \theta_e ... \theta_e + \Delta \theta_e, \varphi ... \varphi + \Delta \varphi)$

Limit of infinitesimal acceptance:

 $\Delta \sigma = \frac{d^3 \sigma}{dE' d\theta_e d\varphi} (E', \theta_e, \varphi) \Delta E' \Delta \theta_e \Delta \varphi = \frac{d\sigma}{dE' d\Omega} \Delta E' \Delta \Omega$ (use Jacobian to transform variables)

In case of more particles/observables and finite phase space:

$$\Delta \sigma = \iiint_{\substack{Phase \\ Space}} \frac{d^n \sigma}{dk_1 dk_2 \dots dk_n} (k_1 \dots k_n) Acc(k_1 \dots k_n) dk_1 dk_2 \dots dk_n$$

where Acceptance is defined by detector boundaries and bin sizes as well as number of observed particles (inclusive = only 1, semi-inclusive, exclusive)

Election Scattering - Theorist's View Target)) (1 center) What is the transition rate Winf? $N_{e,f} = N_{e,in} \cdot P(i \rightarrow f) = I_{e,in} \cdot \frac{N_T}{A} \cdot \Delta \sigma$ = Iein NT . 45 = (Jein) NT . 45 => Wint = Jin . 15 Fermi's GOLDEN Rule: Phase space Wing = 21 Ingil 200 & spanned by detector/kinsmakichin MG: = <45 | Hint | 4in >