

# Lecture 15-16

## 1) Summary QM of particle in 1D

→ state vector  $|\psi\rangle$  represented by  $\psi(x)$  ( $-\infty < x < \infty$ )  
 at every time  $t$   $|\psi\rangle(t) \rightarrow \psi(x,t)$

Normalized:  $\langle \psi | \psi \rangle = \int_{-\infty}^{\infty} \psi^*(x) \psi(x) dx = 1$

a) Evolution in time:  $i\hbar \frac{\partial}{\partial t} |\psi\rangle(t) = H |\psi\rangle(t)$   $H = \frac{p^2}{2m} + V$

$$i\hbar \frac{\partial}{\partial t} \psi(x,t) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi(x,t) + V(x) \psi(x,t)$$

"stationary solution":  $|\psi\rangle_E(t) = E$  to  $H$  with  $H|\psi\rangle_E = E|\psi\rangle_E$   
 $\Rightarrow$  true for all times  $t$ !

$\Rightarrow$  time evolution:  $|\psi\rangle(t) = |\psi\rangle(0) \cdot e^{-i \frac{E}{\hbar} t}$   
 $t=0: \frac{-\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi_E(x) + V(x) \psi_E(x) = E \psi_E(x)$

General solution:  $\psi(x,t) = \sum_E \psi_E(x) e^{-i \frac{E}{\hbar} t}$

b) Probability to find particle at  $x \dots x + \Delta x$  ( $\Delta x$  'small')

$$\Delta P(x \dots x + \Delta x) = |\psi(x,t)|^2 \Delta x$$

In general:

operator  $\Omega \rightarrow$  EV's  $\omega_i$ ; can only find  $\omega_i$  when measure;

collapse  $\rightarrow |\psi\rangle_{\omega_i}$

Probability  $\langle \phi_{\omega_i} | \psi \rangle^2$

Expectation value:  $\langle \Omega \rangle = \langle \psi | \Omega | \psi \rangle$

$\infty$  square well

Ex.:  $\psi_n(x) = \sqrt{\frac{2}{L}} \sin \frac{n\pi x}{L}$

$E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2}$

$P(x \dots x + \Delta x) = \frac{2}{L} \sin^2 \frac{n\pi x}{L} \Delta x$

refer to mid term.

Note:  $\int_0^L |\psi_n(x)|^2 dx = 1!$

2) 3D: generalize  $|\psi\rangle \rightarrow \psi(\vec{r}, t) \quad \vec{r} \in \mathbb{R}^3$

Normaliz.  $\langle \psi | \psi \rangle = \iiint d^3\vec{r} |\psi(\vec{r})|^2 = 1$

a) Hamiltonian:  $\frac{p_x^2 + p_y^2 + p_z^2}{2m} + V(\vec{r}) \rightarrow -\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r})$

still can find separable (stationary) solutions  $\psi_{ES}$   
 $\cdot e^{-i\frac{E}{\hbar}t}$  U/EV E

b)  $\Delta \text{Prob}(\vec{r}, \vec{r} + \Delta \vec{r}) = |\psi(\vec{r})|^2 \Delta \tau$   
Small Volume surrounding  $\vec{r}$

Example: 3D square box  $L^3$ ; combine solutions

$\psi_E(\vec{r}) = \psi_k(x) \cdot \psi_m(y) \cdot \psi_n(z) = \left[ \frac{8}{L^3} \sin \frac{k\pi x}{L} \sin \frac{m\pi y}{L} \sin \frac{n\pi z}{L} \right]$  (point out separable.)

energy  $E_{k,m,n} = \frac{\pi^2 \hbar^2}{2mL^2} (k^2 + m^2 + n^2)$  go through detail!

Note: new phenomenon  $\rightarrow$  degenerate EVs!

eg.  $k=m=n=1$  ground state (only one)

but  $k=2, m=n=1$  or  $k=1, m=2, n=1$  or  $k=1, m=1, n=2$   
 all give same energy! (but different ES!)

c) Example: g.s.  $\rightarrow$  Prob for center? all  $\sin=1 \Rightarrow \frac{8^2}{L^3}$   
 surface edge? 0!

3) What if  $V(\vec{r}) = V(r) \rightarrow$  only depends on distance from origin?

$\rightarrow$  spherical coordinates  $r, \theta, \phi \rightarrow$  EXPLAIN

$\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \rightarrow \frac{1}{r^2} \frac{\partial}{\partial r} r^2 \frac{\partial}{\partial r} + \frac{1}{r^2} \left( \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta} + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} \right)$   
 multiply with  $-\frac{\hbar^2}{2m}$  for H! (2)

3 cont'd: simple "easy" problem: set  $r = \text{const.}$ ,  $\sin \theta = 1$

$\Rightarrow \langle \psi | = \psi(\varphi) \rightarrow \text{solve } -\frac{\partial^2}{\partial \varphi^2} \psi(\varphi) = m^2 \psi(\varphi)$

answ.:  $\psi(\varphi) = e^{\pm i m \varphi}$  but  $\psi(2\pi + \varphi) = \psi(\varphi) \Rightarrow m = \text{int}$

( $0, \pm 1, \pm 2, \dots$ ). Turns out  $\frac{\hbar}{r} \frac{\partial}{\partial \varphi}$  represents  $L_z$

and therefore  $\psi_m(\varphi)$  are ES to  $L_z$  w/ EV  $m\hbar$   
 $\rightarrow$  angular momentum is quantized! (Another one...)

Turns out the whole expression  $-\hbar^2 \nabla^2 = \hat{L}^2$ !

(not obvious) Solutions are simultaneous EV to both

$\hat{L}^2$  (total angular mom.) and  $L_z$ , w/ EV

$\hbar^2 l(l+1)$  and  $\hbar m$ , respectively. They are called

spherical harmonics and are written  $Y_{lm}(\theta, \varphi)$

ES:  $\hat{L}^2 Y_{lm} = \hbar^2 l(l+1) Y_{lm}$   $l = 0, 1, 2, \dots$   
 $L_z Y_{lm} = \hbar m Y_{lm}$  (called "angular momentum")

$m = 0, \pm 1, \dots, \pm l$  (max!)

since  $m$  cannot be  $> l$

Properly normalized:  $Y_{00}(\theta, \varphi) = \frac{1}{\sqrt{4\pi}} = \text{const} = \frac{1}{\sqrt{4\pi}} P_0(\cos \theta)$

(clearly,  $\hat{L}^2 Y_{00} = L_z Y_{00} = 0 \Rightarrow l = m = 0$ )

$Y_{10} = \sqrt{\frac{3}{4\pi}} \cos \theta = \sqrt{\frac{3}{4\pi}} P_1(\cos \theta)$   $Y_{11} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{i m \varphi}$   $l = m = 1$

$Y_{1-1} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{-i m \varphi}$   $Y_{20} = \sqrt{\frac{5}{16\pi}} (3 \cos^2 \theta - 1) = \sqrt{\frac{5}{4\pi}} P_2(\cos \theta)$  etc.